## ORGAN Phase 1B: Results and Future Plans





#### Aaron Quiskamp

![](_page_0_Picture_4.jpeg)

Australian Research Council Centre of Excellence for Engineered Quantum Systems

![](_page_0_Picture_6.jpeg)

![](_page_0_Picture_7.jpeg)

![](_page_1_Picture_0.jpeg)

![](_page_1_Picture_1.jpeg)

## Axions

• Axions can solve two very big problems in physics!

![](_page_2_Picture_2.jpeg)

![](_page_2_Picture_4.jpeg)

## Axions

- Axions can solve two very big problems in physics!
- The strong CP problem and the dark matter problem

![](_page_3_Picture_3.jpeg)

![](_page_3_Picture_6.jpeg)

## Axions

- Axions can solve two very big problems in physics!
- The strong CP problem and the dark matter problem
- Axions may interact with a strong **B** field to produce a photon with frequency related to  $m_a$

![](_page_4_Picture_4.jpeg)

![](_page_4_Picture_5.jpeg)

Inverse Primakoff effect

![](_page_4_Picture_10.jpeg)

![](_page_5_Picture_1.jpeg)

• 
$$f = \frac{m_a}{h}c^2 + \frac{1}{2}\frac{m_a}{h}v^2$$

![](_page_6_Picture_3.jpeg)

• 
$$f = \frac{m_a}{h}c^2 + \frac{1}{2}\frac{m_a}{h}v^2$$

![](_page_7_Figure_3.jpeg)

![](_page_7_Picture_4.jpeg)

• 
$$f = \frac{m_a}{h}c^2 + \frac{1}{2}\frac{m_a}{h}v^2$$

![](_page_8_Figure_3.jpeg)

![](_page_8_Picture_4.jpeg)

![](_page_8_Figure_5.jpeg)

• 
$$f = \frac{m_a}{h}c^2 + \frac{1}{2}\frac{m_a}{h}v^2$$

![](_page_9_Figure_3.jpeg)

![](_page_9_Picture_4.jpeg)

![](_page_9_Figure_5.jpeg)

• 
$$f = \frac{m_a}{h}c^2 + \frac{1}{2}\frac{m_a}{h}v^2$$

![](_page_10_Figure_3.jpeg)

![](_page_10_Picture_4.jpeg)

• Axion mass  $m_a$  determines the real photon frequency (f)

• 
$$f = \frac{m_a}{h}c^2 + \frac{1}{2}\frac{m_a}{h}v^2$$

•  $g_{a\gamma\gamma}$ : Axion-photon coupling strength

![](_page_11_Figure_4.jpeg)

![](_page_11_Picture_5.jpeg)

• 
$$f = \frac{m_a}{h}c^2 + \frac{1}{2}\frac{m_a}{h}v^2$$

- $g_{a\gamma\gamma}$ : Axion-photon coupling strength
- $B_0$ : Magnetic field strength

![](_page_12_Picture_5.jpeg)

![](_page_12_Picture_6.jpeg)

• 
$$f = \frac{m_a}{h}c^2 + \frac{1}{2}\frac{m_a}{h}v^2$$

- $g_{a\gamma\gamma}$ : Axion-photon coupling strength
- $B_0$ : Magnetic field strength
- V: Cavity volume

![](_page_13_Picture_6.jpeg)

![](_page_13_Picture_7.jpeg)

• 
$$f = \frac{m_a}{h}c^2 + \frac{1}{2}\frac{m_a}{h}v^2$$

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- C: Form factor  $\propto \vec{E}_{cav} \cdot \vec{B}_{ext}$

![](_page_14_Picture_7.jpeg)

![](_page_14_Picture_8.jpeg)

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- $Q_L$ : Cavity quality factor

![](_page_15_Picture_8.jpeg)

![](_page_15_Picture_9.jpeg)

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- $Q_L$ : Cavity quality factor
- $\beta$ : Antenna coupling

![](_page_16_Picture_9.jpeg)

![](_page_16_Picture_10.jpeg)

![](_page_17_Picture_1.jpeg)

•  $P_{\text{signal}} \approx \mathcal{O}(10^{-26}) W$ 

![](_page_18_Picture_2.jpeg)

- $P_{\text{signal}} \approx \mathcal{O}(10^{-26}) W$
- Resolve this tiny signal above the noise of our experiment •

![](_page_19_Picture_3.jpeg)

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- Resolve this tiny signal above the noise of our experiment •

• SNR = 
$$\frac{P_{\text{signal}}}{k_{\text{B}}T_{\text{S}}} \sqrt{\frac{\tau}{\Delta v_{\text{a}}}}$$

![](_page_20_Picture_4.jpeg)

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• SNR = 
$$\frac{P_{\text{signal}}}{k_{\text{B}}T_{\text{S}}} \sqrt{\frac{\tau}{\Delta v_{\text{a}}}}$$

<u>Scan rate</u> - How fast we can exclude axions at a given mass and coupling

![](_page_21_Picture_5.jpeg)

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- Resolve this tiny signal above the noise of our experiment lacksquare

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<u>Scan rate</u> - How fast we can exclude axions at a given mass and coupling

$$\frac{df}{dt} = \left(\frac{g_{a\gamma\gamma}^4 \rho_a^2 Q_a}{m_a^2 k_B^2}\right) \frac{B_0^4}{\mathrm{SNR}^2 T_S^2} C^2 V^2 Q_L \frac{\beta}{(1+1)^2}$$

![](_page_22_Picture_6.jpeg)

2

 $-\beta)^2$ 

- $P_{\text{signal}} \approx \mathcal{O}(10^{-26}) W$
- Resolve this tiny signal above the noise of our experiment ullet

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Determined by nature

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Determined by nature Experiment dependent Cavity

![](_page_25_Picture_8.jpeg)

dependent

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Determined by nature Experiment dependent Cavity

![](_page_26_Figure_8.jpeg)

![](_page_27_Picture_1.jpeg)

• High mass (frequency) axion haloscope hosted at UWA

![](_page_28_Picture_2.jpeg)

- High mass (frequency) axion haloscope hosted at UWA
- Why "high mass" (>40µeV)?

![](_page_29_Picture_3.jpeg)

- High mass (frequency) axion haloscope hosted at UWA
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- The high mass parameter space is largely unexplored with many predicitons..

![](_page_30_Picture_4.jpeg)

![](_page_30_Figure_5.jpeg)

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- **SMASH** model predicts  $50 \le m_a \le 200 \, \mu eV$

![](_page_31_Picture_5.jpeg)

![](_page_31_Figure_6.jpeg)

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![](_page_32_Picture_6.jpeg)

![](_page_32_Figure_7.jpeg)

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![](_page_33_Picture_7.jpeg)

![](_page_33_Figure_8.jpeg)

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- **ORGAN** is uniquely positioned to search this range

![](_page_34_Picture_8.jpeg)

![](_page_34_Figure_9.jpeg)

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- **ORGAN** is uniquely positioned to search this range
- $\frac{df}{dt} \propto f^{-14/3} \rightarrow$  High frequency (mass) scales poorly

![](_page_35_Picture_9.jpeg)

![](_page_35_Figure_10.jpeg)


Scan between 15-16 GHz



Tuning rod

Scan between 15-16 GHz







Rotation stage

Tuning rod

Scan between 15-16 GHz







 $\overrightarrow{E}_{cav} \bullet \overrightarrow{B}_{ext} \neq 0$ 





Rotation stage

TM<sub>010</sub> mode

Tuning rod

- Scan between 15-16 GHz
- **Tuning:** moving the rod radially perturbs the axion sensitive mode, shifting the frequency







 $\overrightarrow{E}_{cav} \bullet \overrightarrow{B}_{ext} \neq 0$ 





TM<sub>010</sub> mode

Rotation stage





















#### 'Mode Map'







• What do we mean by scanning?





• What do we mean by scanning?



#### Dilution fridge



#### Step motor



#### • What do we mean by scanning?



**ORGAN DAQ** 





#### • What do we mean by scanning?



ORGAN DAQ







#### • What do we mean by scanning?



ORGAN DAQ









**ORGAN DAQ** 

















 Placing limits 'for free' on other dark matter candidates



 Placing limits 'for free' on other dark matter candidates



#### Limits on Dark Photons, Scalars, and Axion-Electromagnetodynamics with The **ORGAN** Experiment

Ben T. McAllister,<sup>1,2,a</sup> Aaron Quiskamp,<sup>1,b</sup> Ciaran A. J. O'Hare,<sup>3</sup> Paul Altin,<sup>4</sup> Eugene N. Ivanov,<sup>1</sup> Maxim Goryachev,<sup>1</sup> and Michael E. Tobar<sup>1, c</sup>

<sup>1</sup>QDM Laboratory, Department of Physics, University of Western Australia,

35 Stirling Highway, Crawley WA 6009, Australia.

<sup>2</sup>Centre for Astrophysics and Supercomputing, Swinburne University of Technology, John St, Hawthorn VIC 3122, Australia

<sup>3</sup>School of Physics, Physics Road, The University of Sydney, NSW 2006 Camperdown, Sydney, Australia

<sup>4</sup>ARC Centre of Excellence For Engineered Quantum Systems,



- Placing limits 'for free' on other dark matter candidates
- Dark photons convert to detectable photons



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- Placing limits 'for free' on other dark matter candidates
- Dark photons convert to detectable photons
- Simple scaling of Axion limits to **Dark Photon limits**
- Scalar dark matter (eg. dilaton) limits can also be placed



#### Limits on Dark Photons, Scalars, and Axion-Electromagnetodynamics with The **ORGAN** Experiment

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• Search between ~26-27 GHz



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- Length scale ~45% smaller than phase 1a





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- High frequency is difficult —> Resonator is <u>necessarily</u> small
- Relative tolerances are much bigger





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- Extremely sensitive to alignment and rod tilt
- Novel high frequency resonator designs are needed!























- Simple idea by Ben McAllister
- New tunable rectangular cavity solves many problems!




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Parameter	Tuning-rod cavity	Rectangular cavity
С	X	





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С	×	
Q	×	





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Parameter	Tuning-rod cavity	Rectangular cavity
С	×	
Q	×	
V		X





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Parameter	Tuning-rod	Rectangular
	cavity	cavity
С	×	
Q	×	
V		×
Mode crossings	×	





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Parameter	Tuning-rod cavity	Rectangular cavity
С	×	
Q	×	
V		×
Mode crossings	×	
Bore utilisation		×





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Parameter	Tuning-rod cavity	Rectangular cavity
С		
Q	×	
V		X
Mode crossings	×	
<b>Bore utilisation</b>		X
Tuning	×	





- Simple idea by Ben McAllister
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Parameter	Tuning-rod	Rectangular
	cavity	cavity
С	×	
Q	×	
V		×
Mode crossings	×	
Bore utilisation		×
Tuning	×	
Scan rate		=/ 🔽







• First search already complete!





- First search already complete!
- No mode crossings in 26-27 GHz target region!







#### Phase 1

 Targeted searches between 15-16 GHz and 26-27 GHz
~ month scale





#### Phase 2a

- Wider search (15-20 GHz) building on current expertise
  ~ year scale
- Move to mK temperatures and Standard Quantum Limited (SQL) ampifiers





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#### Phase 2a

- Develop efficent single photon counting (SPC) devices
- Reach QCD axion model bands







#### Frequency (GHz)

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- Reach QCD axion model bands









#### (A quick detour..)



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•  $Q \rightarrow Quantum$ 



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- Operates at mK:  $\downarrow T_s$



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•  $Q \rightarrow Quantum$ 



- Utilises a Joshephson Parametric Amplifier (JPA):  $\downarrow T_s$
- Operates at mK:  $\downarrow T_s$

• Variable coupling:  $\uparrow Q_{\rm L} \frac{\beta^2}{(1+\beta)^2}$ 



(A quick detour..)

•  $Q \rightarrow Quantum$ 



- Utilises a Joshephson Parametric Amplifier (JPA):  $\downarrow T_s$
- Operates at mK:  $\downarrow T_s$



• Plan for 5-10 x KSVZ sensitivity





• JPA has optimal gain between 6.1 - 6.4 GHz



- JPA has optimal gain between 6.1 6.4 GHz
- Optimise the cavity for this region → no mode crossings







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- Optimise the cavity for this region → no mode crossings
- Tuning well at mK









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- Optimise the cavity for this region  $\rightarrow$  no mode crossings
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- Final JPA calibrations happening now... ullet









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- JPA has optimal gain between 6.1 6.4 GHz
- Optimise the cavity for this region → no mode crossings
- Tuning well at mK \_\_\_\_\_
- Final JPA calibrations happening now..
- ~1 month scan planned for December
- Set to be most sensitive limits in region
- Will be the first "High-Res" ORGAN search









# **ORGAN Low Frequency**

- Increased interest in low frequency axion searches (<500 MHz) in recent times
- Problem: Cavities get HUGE
- Can use re-entrant cavities to circumvent this issue
- "Cake-like" re-entrant cavity for deployment in large MRI magnet bore at Swinburne
- Experiment under construction













Phase 1 complete





- Phase 1 complete
- Most sensitive limits above 15 GHz





- Phase 1 complete
- Most sensitive limits above 15 GHz
- ORGAN-Q commencing soon..




- Phase 1 complete
- Most sensitive limits above 15 GHz
- ORGAN-Q commencing soon..
- Phase 2 R&D ongoing





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- Most sensitive limits above 15 GHz
- ORGAN-Q commencing soon..
- Phase 2 R&D ongoing
  - Superconducting cavities: 1Q





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  - Superconducting cavities: 1Q
  - Single-photon counting:  $\downarrow T_s$





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- Most sensitive limits above 15 GHz
- ORGAN-Q commencing soon..
- Phase 2 R&D ongoing
  - Superconducting cavities: 1Q
  - Single-photon counting:  $\downarrow T_s$
  - Multiple cavity array: 1V





