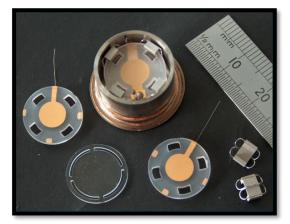


Precision Metrology for Dark Matter Detection



Will Campbell

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Australian Government Australian Research Council





Australian National University









AUSTRALIA



Outline

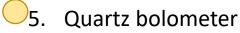
1. Ultralight dark matter, models and motivation

2. Precision metrology



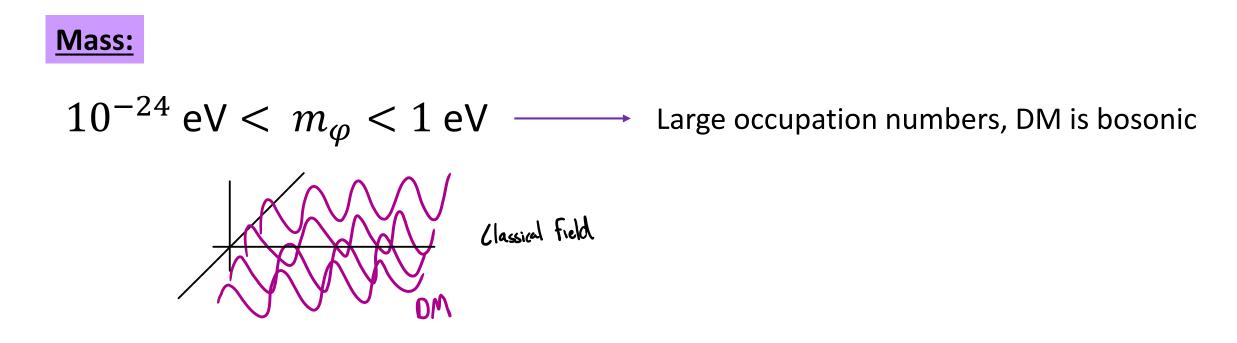
◯3. Quartz bulk acoustic wave resonators, technology and dark matter

4. MAGE





6. Other materials, TeO2 and LiNbO3 BAWs.



Mass:

$$10^{-24}~{
m eV} < m_{arphi} < 1~{
m eV} \longrightarrow$$
 Large occupation numbers, DM is bosonic

Models:

QCD Axion, ALPs, Light Scalars, Vector DM, Fuzzy DM

Mass:

$$10^{-24}~{
m eV} < m_{arphi} < 1~{
m eV} \longrightarrow$$
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Models:

QCD Axion, ALPs, Light Scalars, Vector DM

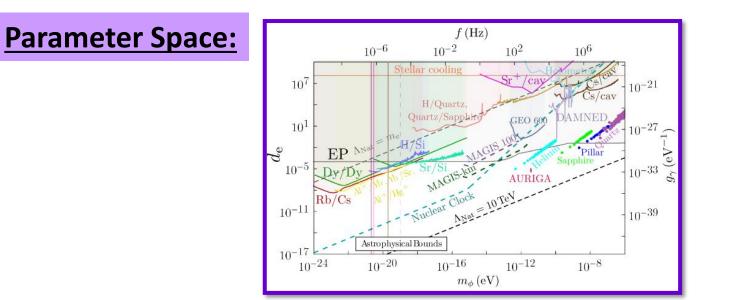
Parameter Space:

Mass:

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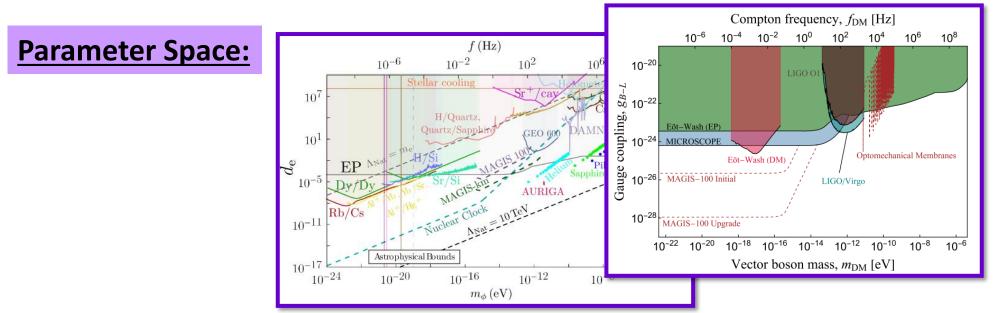


Mass:

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Models:

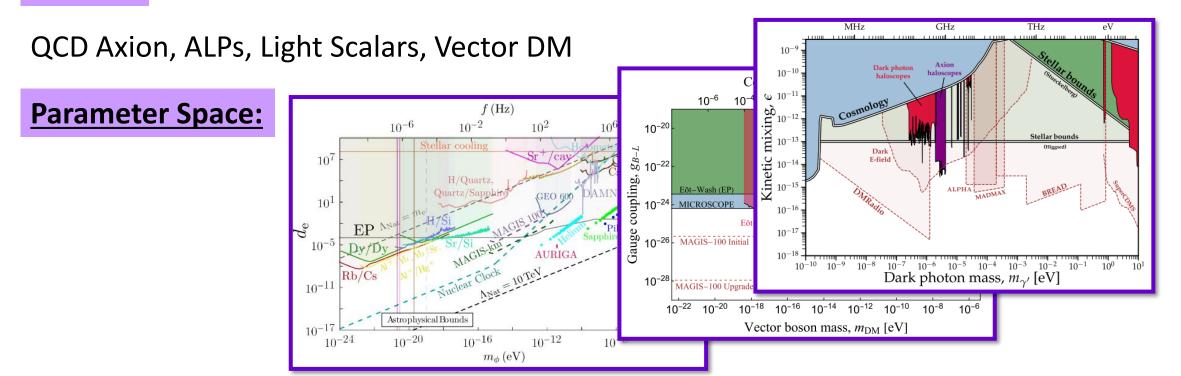
QCD Axion, ALPs, Light Scalars, Vector DM



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Models:





$$10^{-24}~{
m eV} < ~m_{arphi} < 1~{
m eV} ~\longrightarrow~$$
 Large occupation numbers, DM is bosonic

10-

10-2

10-

CROWS

ABRA

10 ct

ALPS-

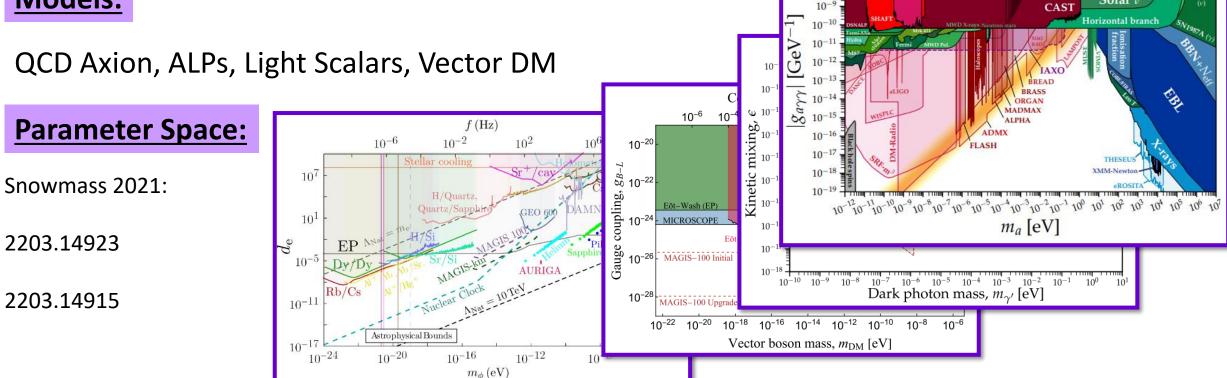
OSOAI

White

SN1987

Solar v

Models:

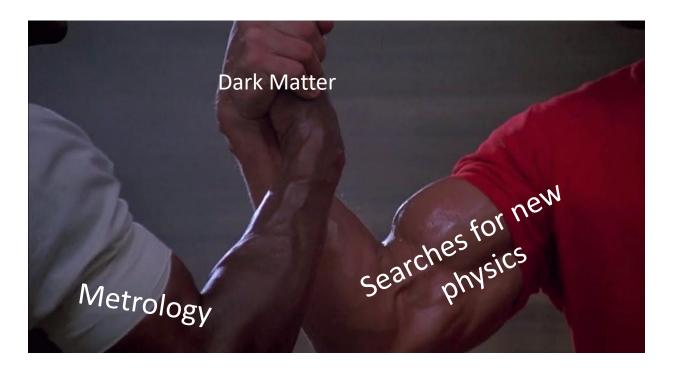


<u>Science of precise measurement</u> — Physics at low energies

<u>Science of precise measurement</u> — Physics at low energies

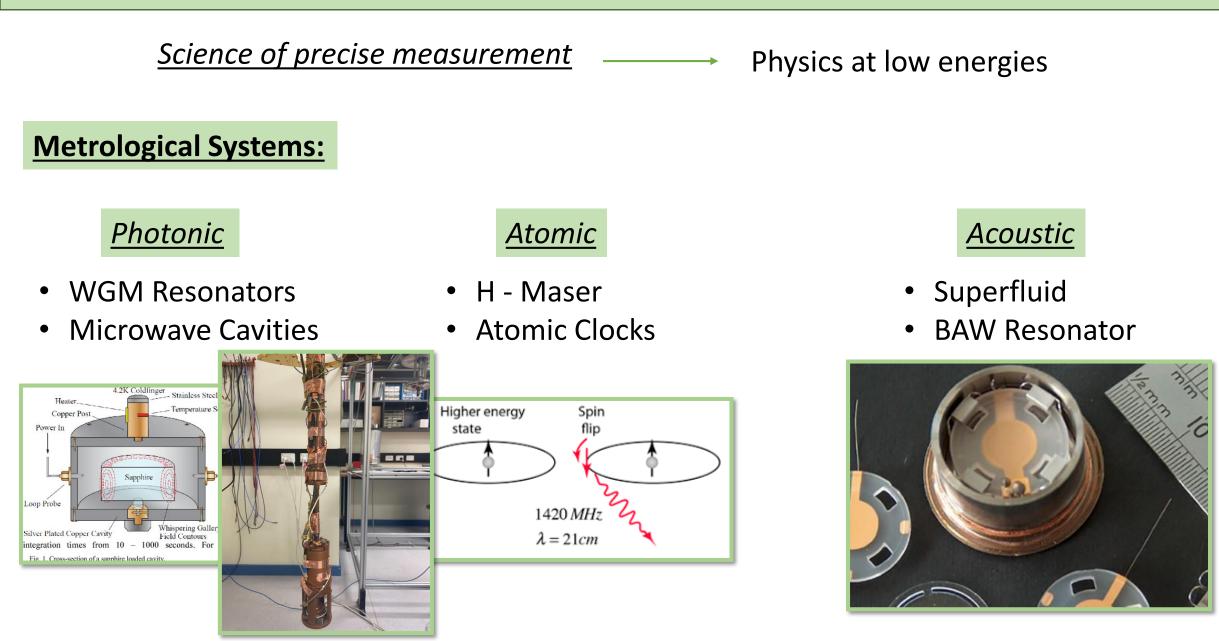


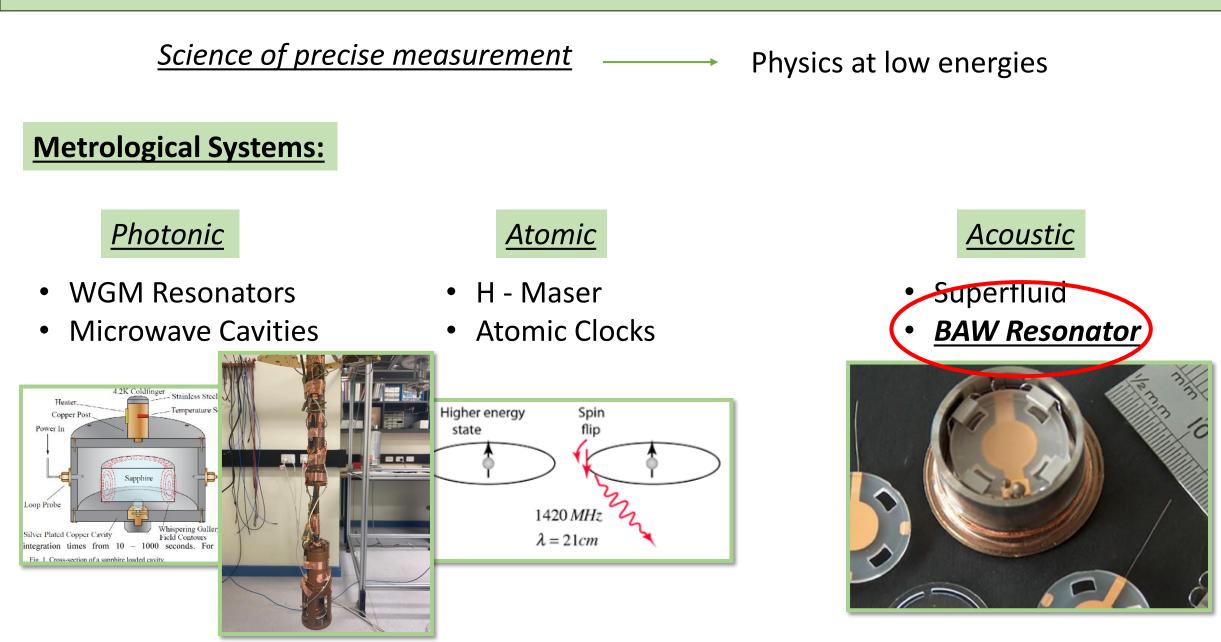
<u>Science of precise measurement</u> — Physics at low energies



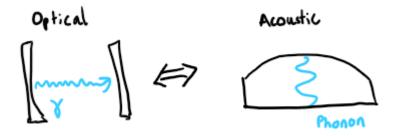
Metrology helps us search for physics beyond the standard model

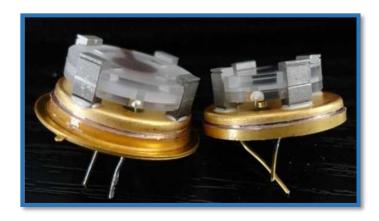
Lorentz violation, fundamental constant variation, tests of general relativity & gravitation, violations of quantum statistics + more

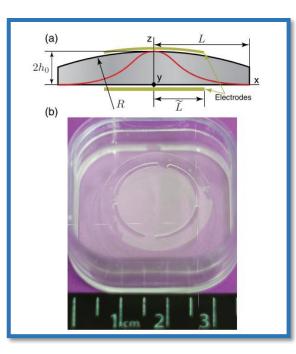




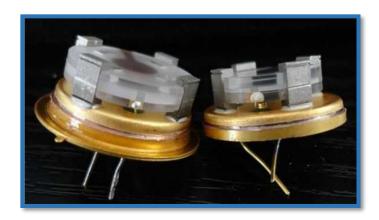
• Acoustic analogue to a Optical Fabry-Perot cavity.

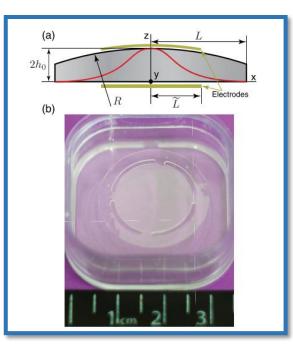




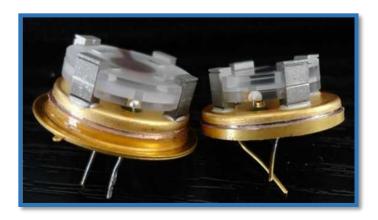


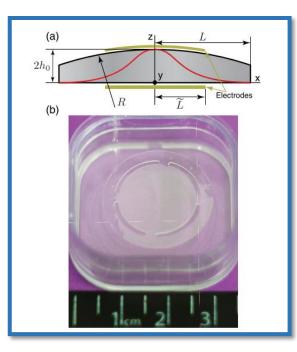
- Acoustic analogue to a Optical Fabry-Perot cavity.
- Already a well established technology



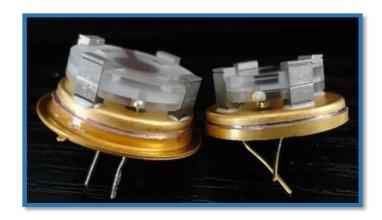


- Acoustic analogue to a Optical Fabry-Perot cavity.
- Already a well established technology
- Gram scale mode mass, macroscopic resonator

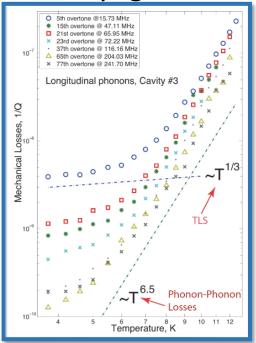


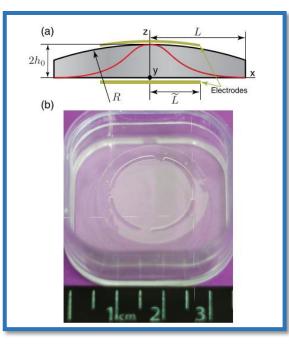


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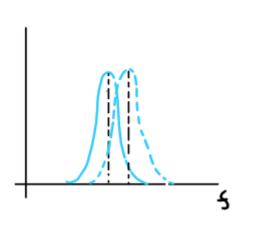
• Extraordinarily high quality factors at cryogenic temperatures $(\sim 10^{10})$

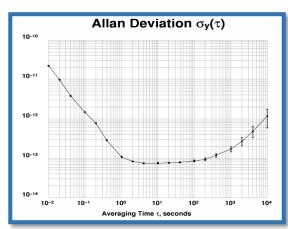


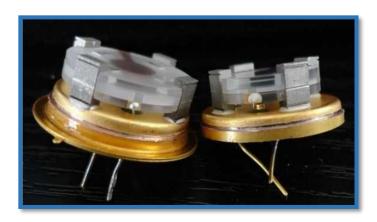


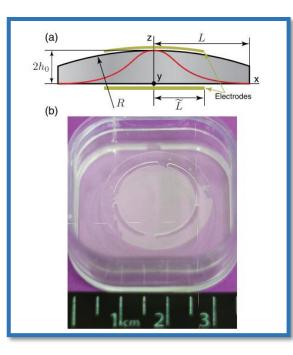
Scientific Reports Vol. 3, 2132 (2013)

- Acoustic analogue to a Optical Fabry-Perot cavity.
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- Impressive short mid term frequency stability



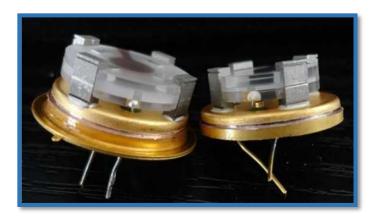


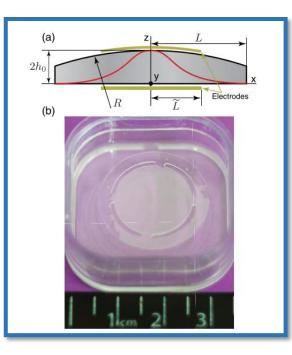




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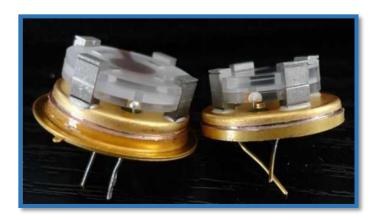


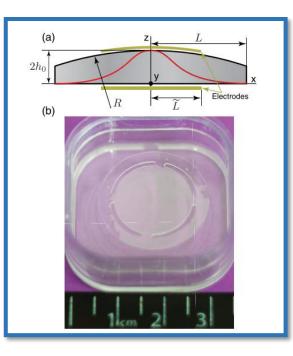




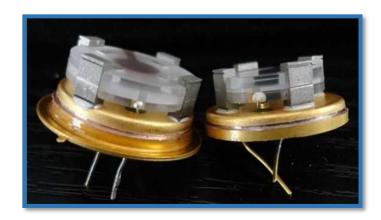
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- High density of modes from 1-1000 MHz

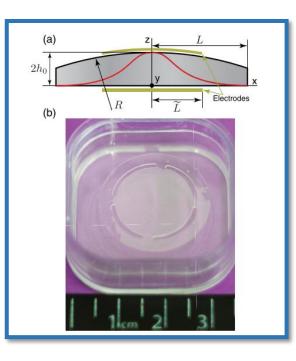






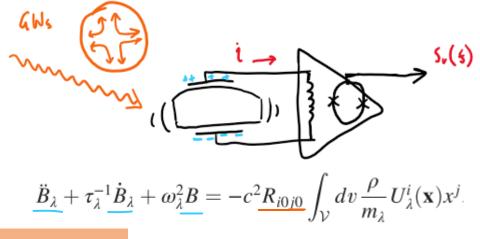
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- Piezoelectric coupling provides excitation & readout
- High density of modes from 1-1000 MHz
- Ongoing studies of behaviour at cryogenic temperatures





Quartz BAW coupled to a DC SQUID amplifier — <u>Highly sensitive resonant mass antenna</u>

Quartz BAW coupled to a DC SQUID amplifier



Primary target:

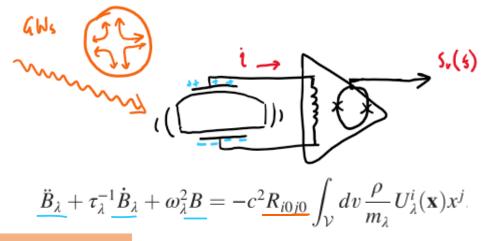
High frequency gravitational waves (MHz)

PRD 90, 102005 (2014)

Highly sensitive resonant mass antenna



Quartz BAW coupled to a DC SQUID amplifier



Primary target:

High frequency gravitational waves (MHz) -

PRD 90, 102005 (2014)

Highly sensitive resonant mass antenna





No known astrophysical sources exist at these frequencies

Quartz BAW coupled to a DC SQUID amplifier

$$\vec{B}_{\lambda} + \tau_{\lambda}^{-1} \vec{B}_{\lambda} + \omega_{\lambda}^{2} \vec{B} = -c^{2} \underline{R}_{i0j0} \int_{\mathcal{V}} dv \frac{\rho}{m_{\lambda}} U_{\lambda}^{i}(\mathbf{x}) x^{j}$$

Primary target:

High frequency gravitational waves (MHz) -

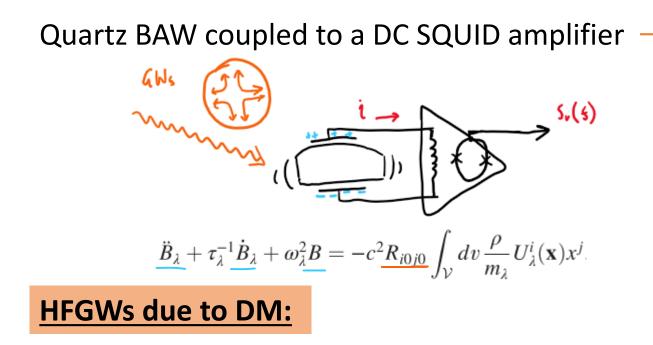
Highly sensitive resonant mass antenna





No known astrophysical sources exist at these frequencies

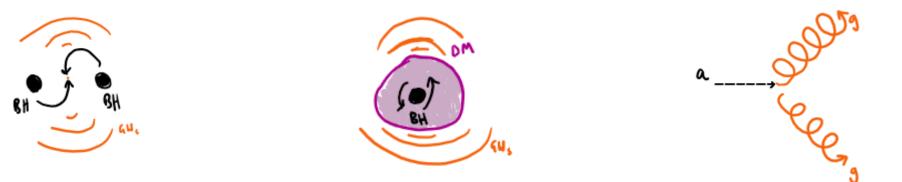
Any potential detection points to new physics outside the standard model !



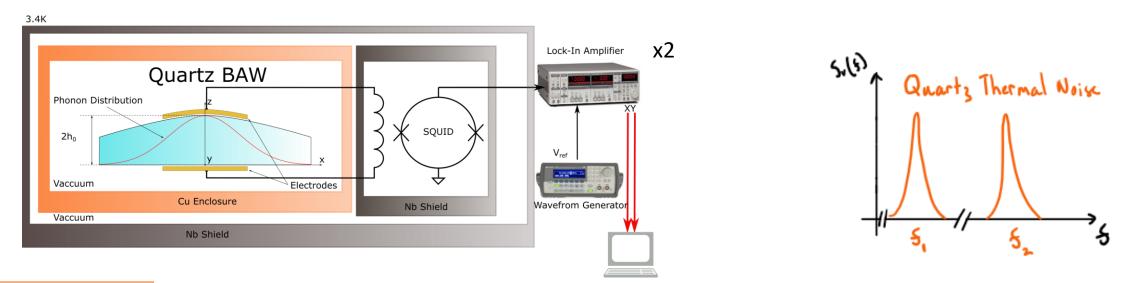
Highly sensitive resonant mass antenna



Sub – solar black hole mergers, black hole super radiance, axion decay into gravitons



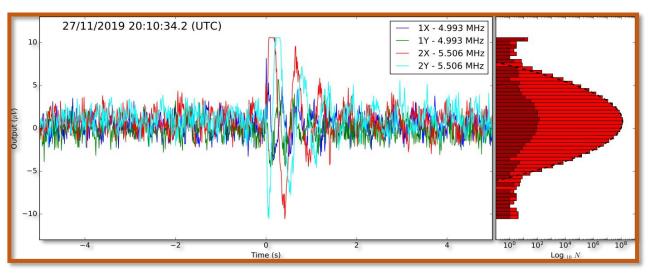
First Observational Period —— GEN 1 & GEN 2, <u>153 days</u> of data, <u>two modes</u>



Data Analysis:

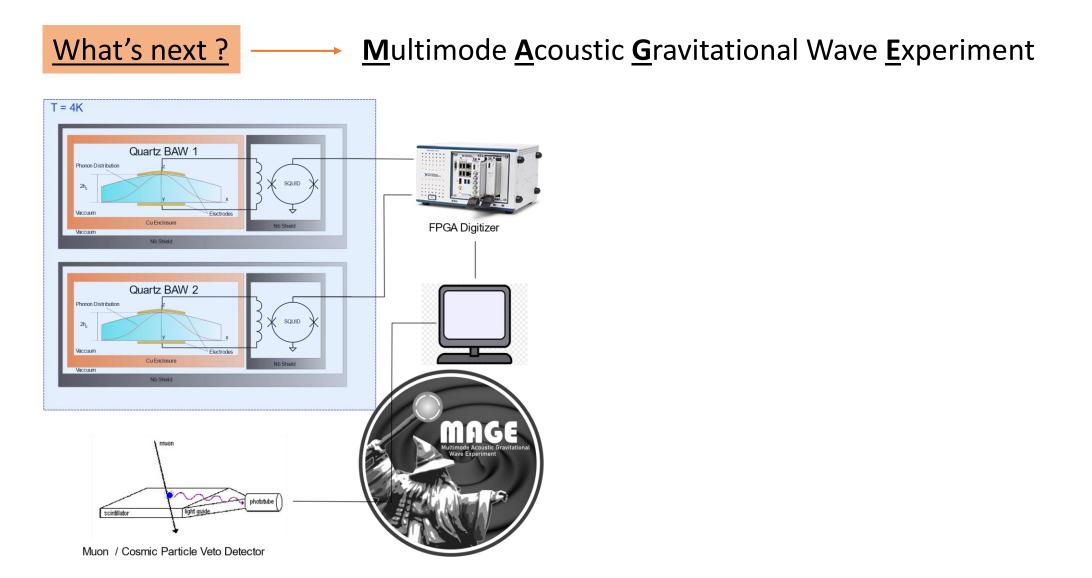
Two significantly strong, rare events

Phys. Rev. Lett. 127, 071102





<u>What's next ?</u> <u>Multimode</u> <u>A</u>coustic <u>G</u>ravitational Wave <u>E</u>xperiment



•

What's next? T = 4KQuartz BAW 1 Phonon Distribution Vaccuu Electrode Cu Enclosure FPGA Digitizer Nb Shield Quartz BAW 2 Phonon Distribution Electrodes Cu Enclosure lilid cintillator Muon / Cosmic Particle Veto Detector

<u>M</u>ultimode <u>A</u>coustic <u>G</u>ravitational Wave <u>E</u>xperiment

• 2 x DC SQUID amplifiers

2 x Quartz BAW crystals

- FPGA DAQ
- Cosmic particle veto (coming soon)

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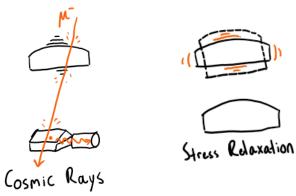
<u>M</u>ultimode <u>A</u>coustic <u>G</u>ravitational Wave <u>E</u>xperiment

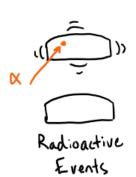
• 2 x DC SQUID amplifiers

2 x Quartz BAW crystals

- FPGA DAQ
- Cosmic particle veto (coming soon)

Exclude potential sources of events:







Currently taking data!

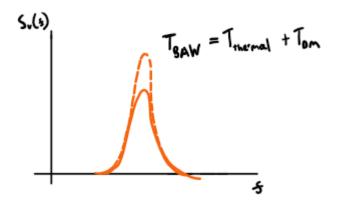


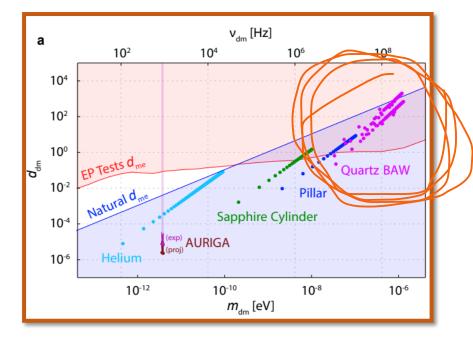
Other possibilities for MAGE:



Other possibilities for MAGE:

Scalar DM -> Isotropic strain signal



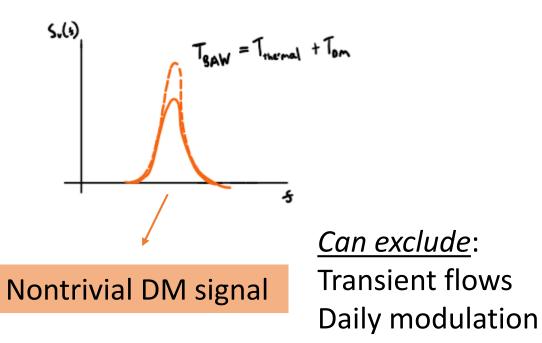


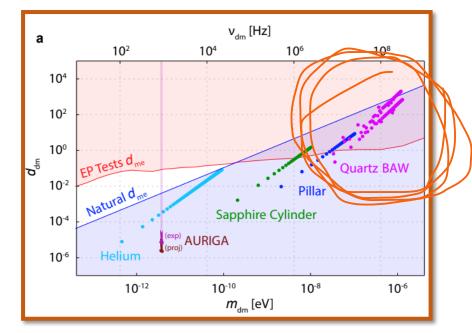
Phys. Rev. Lett. 124, 151301



Other possibilities for MAGE:

Scalar DM -> Isotropic strain signal





Phys. Rev. Lett. 124, 151301

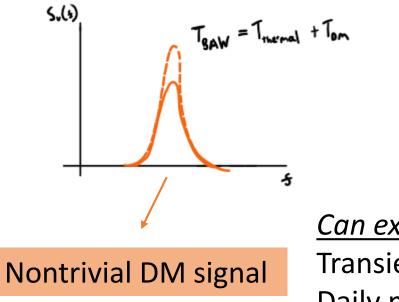
<u>Ongoing work</u>: Resonance tuning



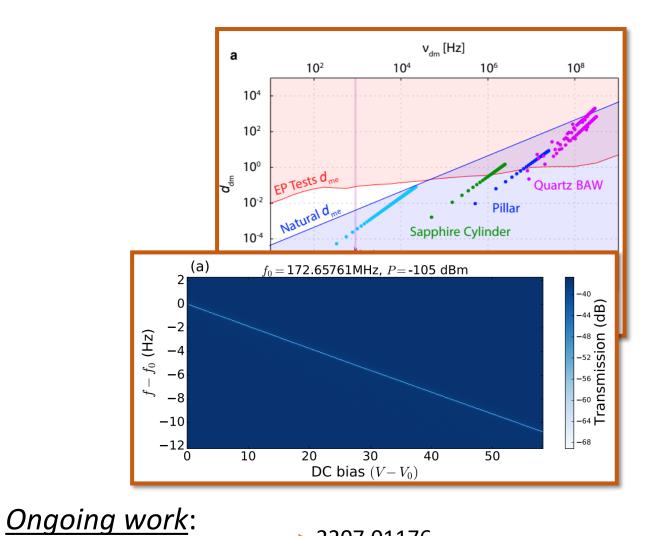
Resonance tuning

Other possibilities for MAGE:

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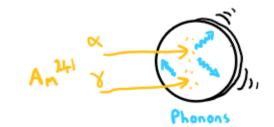
<u>Can exclude</u>: Transient flows Daily modulation



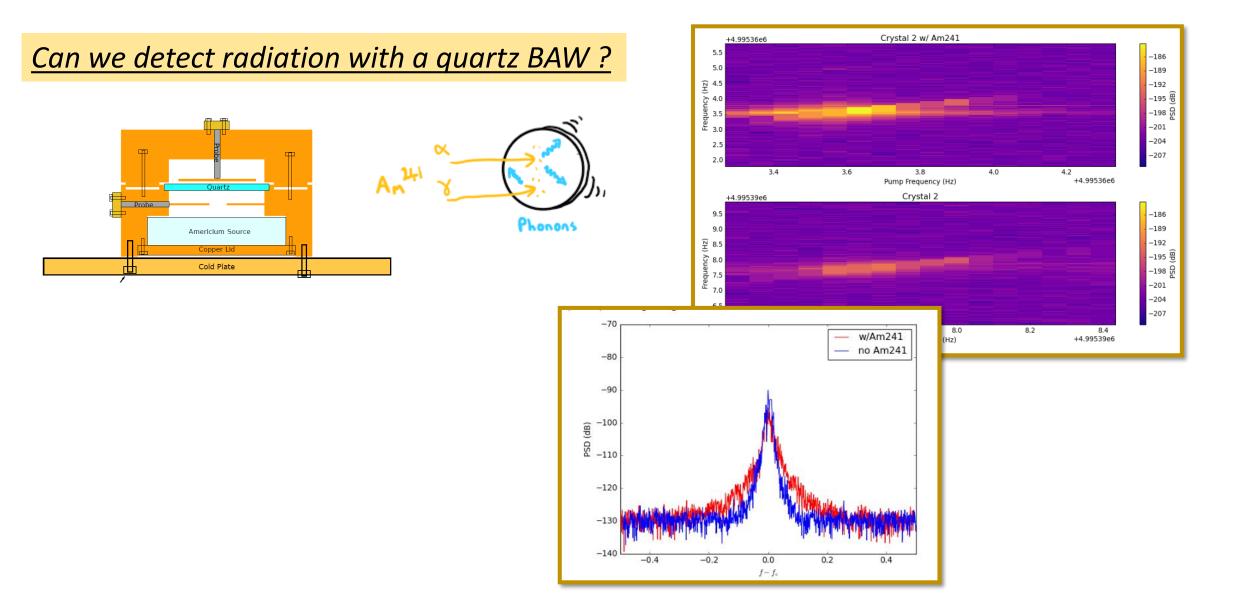
2207.01176

Quartz Bolometer

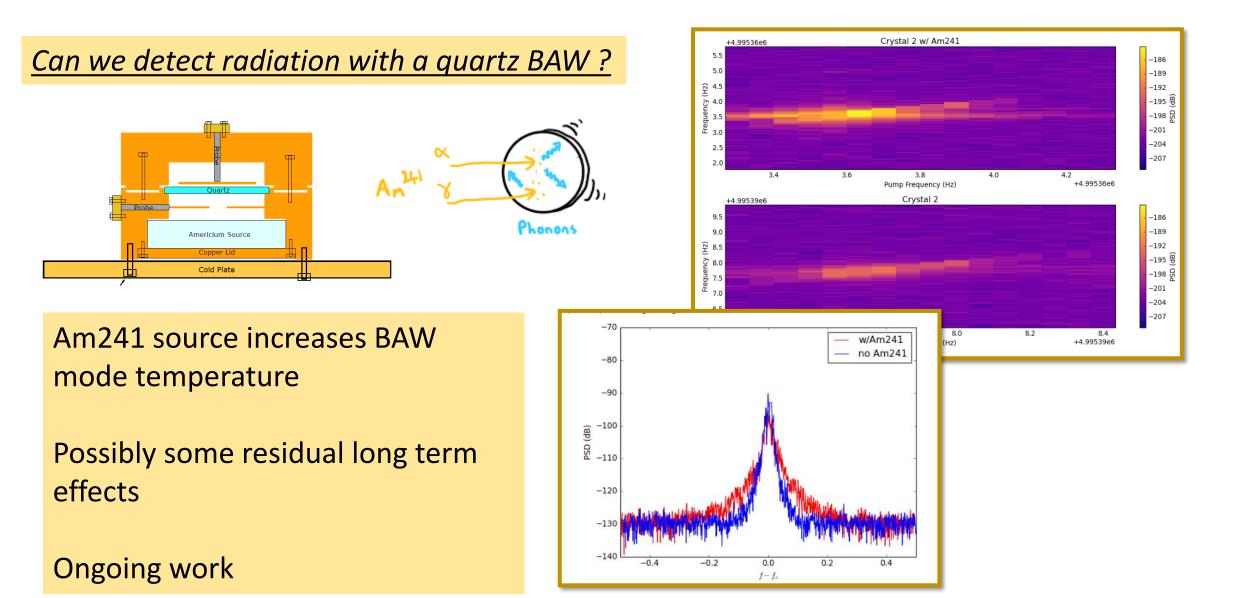
Can we detect radiation with a quartz BAW?



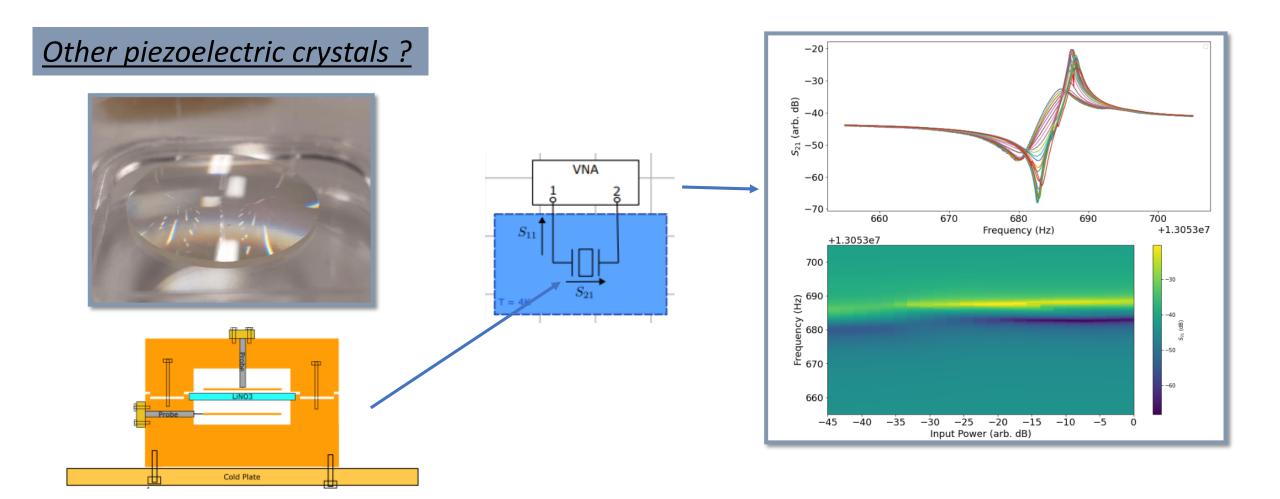
Quartz Bolometer



Quartz Bolometer

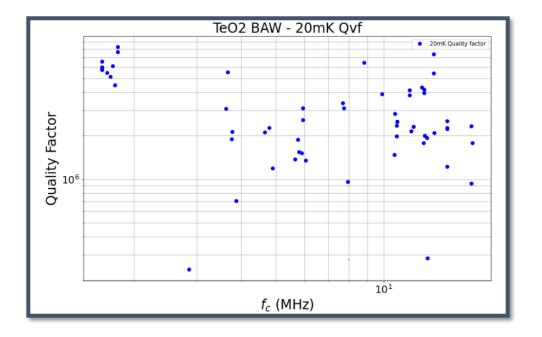


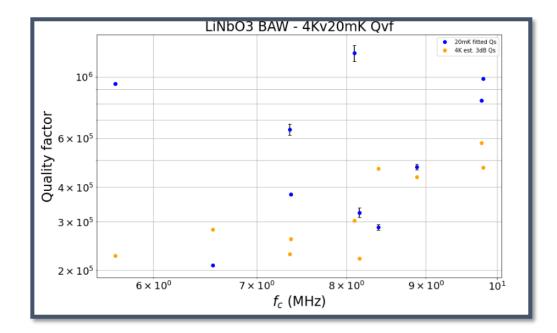
LiNbO3 and TeO2 Bulk Acoustic Wave Resonators



LiNbO3 and TeO2 Bulk Acoustic Wave Resonators

Other piezoelectric crystals ?

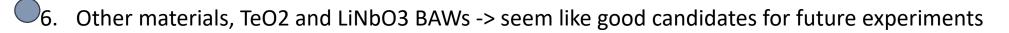




Plenty of modes 1-20 MHz with Quality factors $\sim 10^6$ @ 20mK

Summary

- 1. Many ultralight dark matter models -> Many diverse experiments required
- ●2. Precision metrology provides us with the tools necessary to search for new physics
- O3. Quartz BAWs are an ideal architecture for fundamental physics
- 4. MAGE experiment, overview and outlook
 - 5. Quartz bolometer?









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