

Prototype Particle ID for the SABRE South Liquid Scintillator Veto

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DAMA/LIBRA and Annual Modulation

 If the galaxy sits within a dark matter halo, it is predicted that the rate of direct detections will vary due to the motion of the Earth in the halo

 $R(E) = R_0(E) + R_m \cos(\omega(t - t_0))$

DAMA measures annual modulation at 12.9σ • SABRE to verify using same target material Technology upgrade with use of active veto and muon veto Max in June Min in Dec. 2-6 keV Residuals (cpd/kg/keV) DAMA/LIBRA-phase1 (1.04 ton×yr) DAMA/LIBRA-phase2 (1.13 ton×vr) R. Bernabei et al.. 0.04 Nuc. Phys. Atom. 1 vear 0.02 En. -0.02 Observed modulation at 0.01 cpd/kg/keV -0.04 -0.06 8000 4000 5000 6000 7000 Time (day)

The SABRE South Experiment

Key improvements: 1. High purity Nal(Tl) crystals R5912 Veto PMTs 2. Lower energy threshold – highly efficient PMTs directly coupled to crystals 3. Dual hemisphere data – SUPL and LNGS 4. Active background veto Liquid scintillator veto: • 12 kL of Linear Alkyl-Benzene doped with PPO and bis-MSB 18 Hamamatsu R5912 PMTs Steel and Polyethylene Shielding

High QE and low radioactivity Crystal PMTs + Pure Nal crystals

Steel Vessel containing LS, inner walls to be covered in reflective Lumirror

Muon Detectors

The SABRE South Active Veto

Key purpose:

- Optimal veto of K-40 background in crystals (mimic a DM signal)
- Require >85% veto efficiency of K-40 background





Can we study the backgrounds detected? → Background identification Identify neutron vs. gamma backgrounds?



Energy [keVee]

Liquid Scintillator Detectors

Large scale neutrino detectors most commonly use liquid scintillator due to:

- 1. Ability to **separate energy deposits due to neutrons and photons via pulse shape discrimination**
- 2. Sensitivity to low energies (MeV/keV) i.e. high light yield 12 photons/keV

Necessary for optimal reconstruction of $\bar{v} + p \rightarrow n + e^+$



SNO+

780 tonnes of LAB based LS

Dense array of 25 cm PMTs JUNO

20 ktonnes of LAB based LS

Array of 51 cm Hamamatsu PMTs

Similar/same components as SABRE South LS veto



Pulse Shape Discrimination

Differing interaction mechanisms → Differing proportions of light in pulse

- Neutron vs. Gamma (nuclear vs. electronic interactions)

Pulse shape variables typically exploit different amounts of light in tail for different particles



e.g. Charge ratio variables exploit higher proportion of delayed light emission in neutron interactions

$$Q_{ratio} = rac{Q_{delayed}}{Q_{prompt}}$$

Simplified prototype of veto: SABRINA (little SABRE) – study discrimination b/w gammas/neutrons







New multivariate approach developed to combine variables Ensemble of pulse shape variables fed into multivariate discriminator

This study utilises the 4 pulse shape variable fed into a boosted decision tree (BDT)



Skew: third statistical moment of pulse



Kurtosis: fourth statistical moment of pulse



BDT utilised for multivariate discrimination Compare performance of single cut on commonly used variable to BDT performance BDT (neutron): ~90% efficiency, ~10% fake rate Q_{ratio} (neutron): ~60% efficiency, ~35% fake rate

Understanding of BDT performance at over a range energies is also required

Need to be done via simulation w/ digitisation

Also testing effect of lumirror







Summary

SABRE veto purpose built to veto key WIMP mimics

Background identification being explored via studies of particle ID with prototype detector

Discrimination between neutrons and gamma rays with variables exploiting different pulse shapes

BDT combining PSD variables showing promise with improved neutron/gamma discrimination